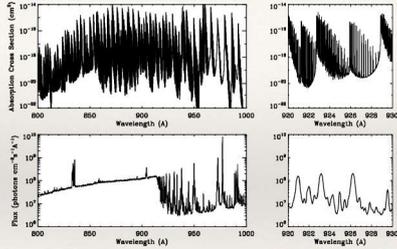


How can we explain all these steps of haze/cloud formation and evolution?

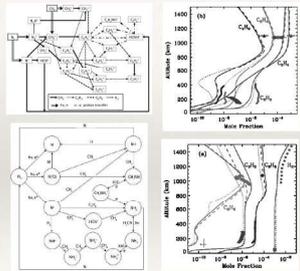
Titan after Cassini-Huygens, 2025, Elsevier
Chapter 7 - Titan's atmospheric structure, composition, haze, and dynamics (Vuitton et al. 2025)

N₂

Photons

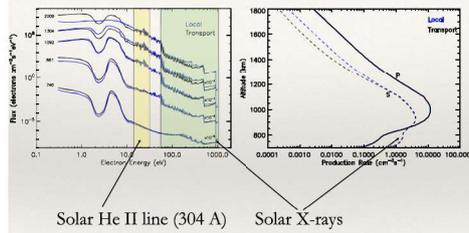


(Walter, Cosby and Helm, 1993)
1021 Å > λ > 891 Å : N(²D) + N(⁴S)
891 Å > λ > 854 Å : N(²P) + N(⁴S)
λ < 854 Å : N(²D) + N(²D)



Models
Srobel et al. 1978
Yang et al. 1984
Toublanc et al. 1995
Lara et al. 1996
Labonville et al. 2001
Wilson & Atreya 2004
Vuitton et al. 2007/8
Horn et al. 2008
Larvas et al. 2008a,b
Kranopolsky 2009/12
Yelle et al. 2010
Larvas et al. 2011
Mandic et al. 2012
Vuitton et al. 2012
Vuitton et al. 2014
Lobson et al. 2015
Vuitton et al. 2019
Lolson et al. 2019
100s - 1000s of reactions
Immense efforts from experimental & theoretical investigations

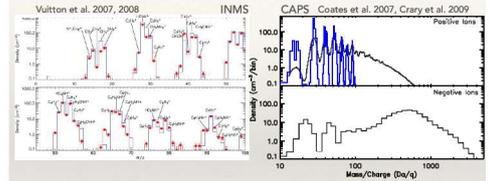
Haze formation



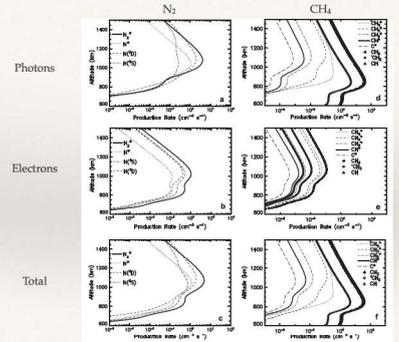
Solar He II line (304 Å) Solar X-rays

Microphysics of photochemical haze formation & evolution

Formation



Gas to particle transition in the ionosphere



MOLECULAR GROWTH

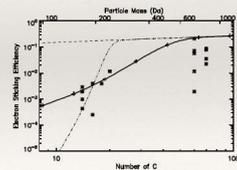
Photochemical products
Positive ions
electrons
Macromolecules from ion chemistry

MOLECULAR GROWTH

Photochemical products
Macromolecules

MOLECULAR GROWTH

Macromolecules = electrons attach on macromolecules



MOLECULAR GROWTH

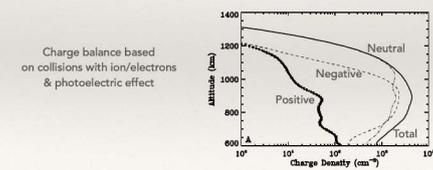
Macromolecules = Charged macromolecules attract positive ions

MOLECULAR GROWTH

Macromolecules = Recombination leads to mass transfer to macromolecules

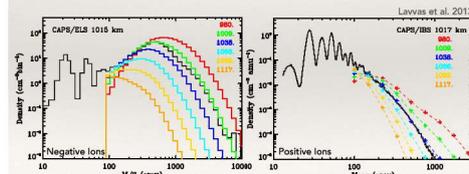
MOLECULAR GROWTH

Macromolecules = Recombination leads to mass transfer to macromolecules



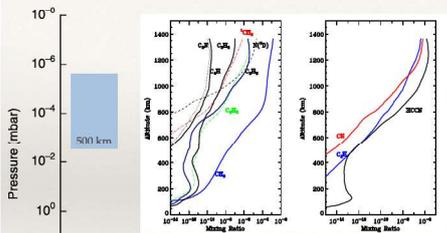
MOLECULAR GROWTH

Macromolecules = Recombination leads to mass transfer to macromolecules



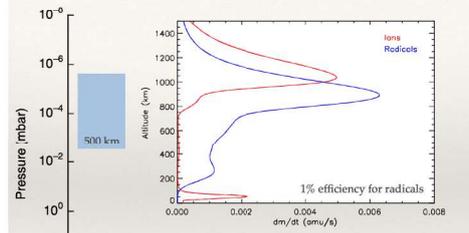
So, ion chemistry is starting the haze formation!
What provides further mass on the particles, though?

Heterogeneous neutral chemistry



Need constraints on heterogeneous reaction efficiency of different species
H on tholins: Sekine et al. 2008a,b

Haze lifetime on Titan: Heterogeneous neutral chemistry



Need constraints on heterogeneous reaction efficiency of different species
H on tholins: Sekine et al. 2008a,b

Adsorption
Desorption
Chemisorption
Ageing

CH ₃	10 ⁻⁴	von Keudell et al. 2002	Exp
CH ₃ H	10 ⁻²		
C ₂	~1	Neyts et al. 2006	MD
C ₂ H	0.80	Hopfi et al. 2000	Exp
C ₂ H ₃	0.35	Hopfi et al. 2000	Exp
C ₂ H ₅	0.001	Hopfi et al. 2000	Exp
C ₃	~0.5	Neyts et al. 2006	MD
I-C ₃ H	0.1-0.3	Neyts et al. 2006	MD
c-C ₃ H	0.97	Neyts et al. 2006	MD

Extended literature in studies on C:H films and molecular dynamic studies.